

# The radiometer equation – a more careful look

In the lectures we derive the equation

$$\text{SNR} = \frac{T_A}{T_{\text{sys}}} (\Delta\nu\tau)^{1/2}, \quad (1)$$

sometimes called the *radiometer equation*, to determine the signal-to-noise ratio from a source generating an antenna temperature  $T_A$  using a telescope with a system temperature  $T_{\text{sys}}$ , a bandwidth  $\Delta\nu$  and an integration time  $\tau$ . Although this result is rigorously correct we made a couple of hand-wavy approximations along the way, so here we'll do it more carefully.

A radio source (and any other contributing noise source) will generate a randomly fluctuating voltage  $u(t)$  at the antenna which we will assume has a gaussian probability distribution with a standard deviation  $\sigma$  and a mean of zero, so that the probability density for  $u$  is

$$p(u) = \frac{1}{(2\pi)^{1/2}\sigma} \exp\left(-\frac{u^2}{2\sigma^2}\right). \quad (2)$$

Our radio telescope system generates a signal  $V(t)$ , proportional to  $[u(t)]^2$ , which we use as our estimator for the noise power. If the signal from the source is the only noise present, then the mean of  $V$  is proportional to the source's flux density. If there are other noise sources present then only part of  $V$  will be from the source. The pdf of  $V$  is one-sided ( $V \geq 0$ ), and is related to  $p(u)$  by

$$p(V) |dV| = 2p(u) |du| \quad (V \geq 0). \quad (3)$$

Using

$$dV = 2u du, \quad (4)$$

we can write

$$p(V) = \frac{V^{-1/2}}{(2\pi)^{1/2}\sigma} \exp\left(-\frac{V}{2\sigma^2}\right). \quad (5)$$

This is known as a *chisquared distribution with one degree of freedom*, and has a mean of

$$\langle V \rangle = \int_0^\infty V p(V) dV = \sigma^2, \quad (6)$$

and a variance of

$$\text{var}[V] = \langle V^2 \rangle - \langle V \rangle^2 \quad (7)$$

$$= \int_0^\infty V^2 p(V) dV - \sigma^4 \quad (8)$$

$$= 2\sigma^4. \quad (9)$$

The standard deviation of  $V$  is therefore  $\sqrt{2}\sigma^2$ , or  $\sqrt{2}\langle V \rangle$ . If the random signal is the only 'noise' in the system, then the signal-to-noise ratio is simply

$$\text{SNR} = \frac{\langle V \rangle}{[\text{var}(V)]^{1/2}} = \frac{1}{\sqrt{2}}. \quad (10)$$

In the lectures we use a plausibility argument to say that this signal-to-noise ratio is ' $\simeq 1$ ', but the result above is more rigorous. If other noise sources are present things are a little more complicated, but the basic argument still holds. The signal will now be represented by the power from just the source, proportional to the antenna temperature  $T_A$ , and the total noise power by the system temperature  $T_{\text{sys}}$ , so that

$$\text{SNR} = \frac{T_A}{T_{\text{sys}}\sqrt{2}}. \quad (11)$$

We can now consider how to improve this signal-to-noise ratio by integrating (averaging) the sample values of  $V$  over some interval of time  $\tau$ . Let there be  $N$  samples of  $V$  (or  $u$ ) in time  $\tau$ . If the samples are statistically independent, and  $N \gg 1$ , then (by definition) the standard deviation of the average is  $\sqrt{N}$  less than that of a sample signal-to-noise ratio, i.e.,

$$\text{SNR} = \frac{T_A}{T_{\text{sys}}\sqrt{2}}\sqrt{N}. \quad (12)$$

Samples  $u_i$  and  $u_j$  will be independent if

$$\langle u_i u_j \rangle = \langle u^2 \rangle \delta_{ij}, \quad (13)$$

i.e., if the autocorrelation of  $u$  is a delta function at the origin. This corresponds (by the Wiener-Khinchin theorem) to a white ('flat') power spectrum for the noise. The Nyquist sampling theorem tells us that a sampling interval of  $\tau/N$  corresponds to a bandwidth of  $\Delta\nu$  where

$$\Delta\nu = \frac{1}{2} \frac{N}{\tau}, \quad (14)$$

so

$$N = 2\Delta\nu\tau, \quad (15)$$

(in the lectures we say that ' $N \simeq \Delta\nu\tau$ '), and therefore our final signal-to-noise ratio is

$$\text{SNR} = \frac{T_A}{T_{\text{sys}}2^{1/2}}(2\Delta\nu\tau)^{1/2} = \frac{T_A}{T_{\text{sys}}} (\Delta\nu\tau)^{1/2}. \quad (16)$$

Note that the two factors of  $\sqrt{2}$  introduced by our more careful analysis cancel out, so the precise and 'approximate' results are the same.